

High Energy Physics – Phenomenology

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## ABSTRACT

We re-consider recent measures of  $R_K$  and  $R_{K^*}$ , now compatible with the Standard Model expectations, as well as the results for the process  $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$  alongside earlier determinations of  $R_{D^{(*)}}$  and  $\text{BR}(B_c \rightarrow \tau \nu)$ . We provide analytic constraints on the associated Wilson coefficients in both the  $b \rightarrow s$  and the  $b \rightarrow c$  sectors. These allow us to estimate the scale of potential New Physics for generic extensions of the Standard Model. We then use the results to constrain the leptoquark landscape.

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## 1. Introduction

Experimental tests at colliders of the Standard Model have crowned it as the golden standard of our understanding of fundamental interactions. However, the model is far from being satisfactory. From a theoretical standpoint, there is no underlying explanation for the gauge structure nor why we observe three matter generations, the Higgs sector is unnatural and the theory, in absence of gravity, develops an UV cutoff (Landau Pole) and there is yet no consensus on a consistent theory of quantum gravity and particle physics. On the experimental front, the Standard Model fails to explain the matter-anti matter asymmetry, it lacks a candidate for dark matter and does not address the origin of neutrino masses. It is therefore fair to say that we have so far established an excellent “Effective Model” of fundamental interactions and that sometime (hopefully) soon even collider experiments will provide clear indications on which extension should be considered as the new Standard Model.

Of course, any extension of the Standard Model will have to reproduce its successes at lower energies, and one can use this to narrow the pathways for new physics models. In our recent review [1] we laid the foundations for a general investigation of models of new physics stemming from experimental test of the Standard Model, from B physics to lepton  $g - 2$ .

Given the experimental updates such as the ones about lepton-flavor universality [2,3] we re-analyze their impact on possible deviations from the Standard Model while also taking into account the anomalies in the  $R_{D^*}$  measurements [4] that we had omitted in [1]. Additionally we consider also the still significant deviations from the SM predictions in rare  $B$  meson decays [5–19]. In particular, we will consider the leptoquark landscape which has been often used in the literature to discuss new physics in the flavour sector either as elementary extensions of the Standard Model or as effective descriptions of some underlying composite dynamics. As we shall see, our estimates can be used to guide searches of new physics at present and future colliders.

The work is organized as follows. In Section 2 and 3 we re-analyse the experimental results and their impact on the associated theoretical effective description in terms of Wilson coefficients for  $b \rightarrow c$  and  $b \rightarrow s$  observables. An analytic study is performed for those observables that are less affected by hadronic physics contamination which are  $R_{K^{(*)}}$ ,  $R_{D^{(*)}}$ ,  $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$  and  $\text{BR}(B_c \rightarrow \tau \nu)$ . We further compare the analytic results with numerical fits using FLAVIO [20] consolidating our results. Having gained information on the maximum deviation allowed by the Wilson coefficients we link them to time-honoured underlying models in Section 4 such as the  $Z'$  and different type of scalar and vector leptoquarks (LQ). Assuming our focus on models of new physics (NP) in which the new couplings are not too finely tuned, we show that to simultaneously accommodate Standard Model deviations in  $R_{D^{(*)}}$ , while respecting the constraints coming from  $R_{K^{(*)}}$ , one can employ the weak singlet scalar leptoquark  $S_1$ , with a mass in the TeV range. Additionally, our analysis suggests that alternative leptoquark models require large fine-tuned couplings to effectively account for both  $b \rightarrow s$  and  $b \rightarrow c$  data. Subsequently, we furnish bounds on the possible NP scales.

## 2. $b \rightarrow s$ flavour constraints

To explore NP emerging around the electroweak energy scale affecting the bottom to strange transitions involving leptons one is led to introduce the following class of effective operators

$$\mathcal{O}_{b_X \ell_Y} = (\bar{s} \gamma_\mu P_X b)(\bar{\ell} \gamma_\mu P_Y \ell), \quad (1)$$

which can be written as  $\text{SU}(2)_L$ -invariant operators. A more general discussion can be found in [21–23]. These operators are incorporated in the following effective Hamiltonian

$$\mathcal{H}_{\text{eff}} = -V_{tb} V_{ts}^* \frac{\alpha_{\text{em}}}{4\pi v^2} \sum_{\ell, X, Y} C_{b_X \ell_Y} \mathcal{O}_{b_X \ell_Y} + \text{h.c.}, \quad (2)$$

where the sum runs over leptons  $\ell = \{e, \mu, \tau\}$  and over their chiralities  $X, Y = \{L, R\}$ . Additionally, it is convenient to define dimensionless coefficients  $C_I$  related to the dimensionful  $c_I$  coefficients appearing in the equivalent Lagrangian formulation

$$\mathcal{L}_{\text{eff}} = \sum_{\ell, X, Y} c_{b_X \ell_Y} \mathcal{O}_{b_X \ell_Y}, \quad \text{with} \quad c_I = V_{tb} V_{ts}^* \frac{\alpha_{\text{em}}}{4\pi v^2} C_I, \quad (3)$$

where  $V_{ts} = -0.0412 \pm 0.0006$  and  $v = (2\sqrt{2}G_F)^{-1/2} = 174$  GeV is the Higgs vacuum expectation value and  $G_F$  the Fermi constant.

### 2.1. Applications to $R_{K^{(*)}}$ and $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$

As main applications of the formalism above we start by considering the well known  $R_K$  and  $R_{K^*}$  ratios [6–8,12]

$$R_K = \frac{\text{BR}(B^+ \rightarrow K^+ \mu^+ \mu^-)}{\text{BR}(B^+ \rightarrow K^+ e^+ e^-)}, \quad R_{K^*} = \frac{\text{BR}(B \rightarrow K^* \mu^+ \mu^-)}{\text{BR}(B \rightarrow K^* e^+ e^-)}. \quad (4)$$

These quantities are known to be excellent tests of lepton flavour universality since they are constructed to reduce QCD-related uncertainties [24]. Therefore, together with the  $B_s \rightarrow \mu^+ \mu^-$  branching ratio [9,13,14] are hadronic insensitive [25] quantities. Here, ‘hadronic insensitive’ (HI) [25] refers to the fact that the related observables have at most few percent QCD-induced theoretical errors.

The latest experimental results for  $R_{K^{(*)}}$  read [2,3]:

$$R_K = 0.994_{-0.082}^{+0.090}(\text{stat})_{-0.027}^{+0.029}(\text{syst}) \quad \text{with } q^2 \in [0.1, 1.1] \text{ GeV}^2, \quad (5)$$

$$R_{K^*} = 0.927_{-0.087}^{+0.093}(\text{stat})_{-0.035}^{+0.036}(\text{syst}) \quad \text{with } q^2 \in [0.1, 1.1] \text{ GeV}^2, \quad (6)$$

$$R_K = 0.949_{-0.041}^{+0.042}(\text{stat})_{-0.022}^{+0.022}(\text{syst}) \quad \text{with } q^2 \in [1.1, 6] \text{ GeV}^2, \quad (7)$$

$$R_{K^*} = 1.027_{-0.068}^{+0.072}(\text{stat})_{-0.026}^{+0.027}(\text{syst}) \quad \text{with } q^2 \in [1.1, 6] \text{ GeV}^2. \quad (8)$$

At high  $q^2$  one can neglect the lepton masses and the formulae for  $R_K$  and  $R_{K^*}$  simplify to<sup>1</sup>

$$R_K = \frac{|C_{b_{L+R}\mu_{L-R}}|^2 + |C_{b_{L+R}\mu_{L+R}}|^2}{|C_{b_{L+R}e_{L-R}}|^2 + |C_{b_{L+R}e_{L+R}}|^2}, \quad (9)$$

$$R_{K^*} = \frac{(1-p)(|C_{b_{L+R}\mu_{L-R}}|^2 + |C_{b_{L+R}\mu_{L+R}}|^2) + p(|C_{b_{L-R}\mu_{L-R}}|^2 + |C_{b_{L-R}\mu_{L+R}}|^2)}{(1-p)(|C_{b_{L+R}e_{L-R}}|^2 + |C_{b_{L+R}e_{L+R}}|^2) + p(|C_{b_{L-R}e_{L-R}}|^2 + |C_{b_{L-R}e_{L+R}}|^2)}, \quad (10)$$

where  $p \approx 0.86$  is the ‘polarisation fraction’ [26–28], that is defined as

$$p = \frac{g_0 + g_{\parallel}}{g_0 + g_{\parallel} + g_{\perp}}. \quad (11)$$

The  $g_i$  are the contributions to the decay rate (integrated over the intermediate bin) of the different helicities of the  $K^*$ . The index  $i$  distinguishes the various helicities: longitudinal ( $i=0$ ), parallel ( $i=\parallel$ ) and perpendicular ( $i=\perp$ ).

We also use the notation [1,25]:

$$C_{b_{L\pm R}\ell_Y} \equiv C_{b_L\ell_Y} \pm C_{b_R\ell_Y}, \quad C_{b_{L+R}\ell_{L\pm R}} \equiv C_{b_L\ell_L} + C_{b_R\ell_L} \pm C_{b_L\ell_R} \pm C_{b_R\ell_R}. \quad (12)$$

The coefficients  $C_{b_X\ell_Y}$  are linked to  $C_9$  and  $C_{10}$  via

$$2C_9 = C_{b_L\ell_{L+R}}, \quad 2C_{10} = -C_{b_L\ell_{L-R}}, \quad 2C'_9 = C_{b_R\ell_{L+R}}, \quad 2C'_{10} = -C_{b_R\ell_{L-R}}. \quad (13)$$

We split the coefficients into the sum of the Standard Model contribution and the beyond Standard Model one via

$$C_{b_X\ell_Y} = C_{b_X\ell_Y}^{SM} + C_{b_X\ell_Y}^{BSM}. \quad (14)$$

We use the Standard Model values given in [1,25] to numerically specify expressions (9)-(10):  $C_{b_L\ell_L}^{SM} = 8.64$  and  $C_{b_L\ell_R}^{SM} = -0.18$ , while  $C_{b_R\ell_X}^{SM} = 0$  for  $X = L, R$ .

We assume that possible NP corrections occur in the muonic Wilson coefficients (i.e.  $C_{b_X\ell_Y}^{BSM} = 0$ ), implying Lepton Flavor Universality Violation (LFUV). Of course, NP effects could appear in other sectors as well (see e.g. [29,30]). However, experimentally muons are easier to monitor and therefore we focus on this possibility. Fig. 1 is the update of Fig. 3 in [1,25] and describes the effects on  $R_{K^*}$  and  $R_K$  obtained switching on one NP Wilson coefficient at a time. The recent experimental measurements are exemplified at one and two sigma by the green  $q^2$  bin  $[1.1, 6] \text{ GeV}^2$  and dark-grey  $q^2$  bin  $[0.1, 1.1] \text{ GeV}^2$  crosses. Substituting in (9)-(10) the SM values for the Wilson coefficients, we obtain

$$\begin{aligned} R_K &= 1 + 0.231[\text{Re}(C_{b_L\mu_L}^{BSM}) + \text{Re}(C_{b_R\mu_L}^{BSM})] - 0.005[\text{Re}(C_{b_L\mu_R}^{BSM}) + \text{Re}(C_{b_R\mu_R}^{BSM})] + \\ &\quad + 0.013\{[\text{Re}(C_{b_L\mu_L}^{BSM}) + \text{Re}(C_{b_R\mu_L}^{BSM})]^2 + [\text{Re}(C_{b_L\mu_R}^{BSM}) + \text{Re}(C_{b_R\mu_R}^{BSM})]^2\}, \\ R_{K^*} - R_K &= -0.398\text{Re}(C_{b_R\mu_L}^{BSM}) + 0.008\text{Re}(C_{b_R\mu_R}^{BSM}) \\ &\quad - 0.046[\text{Re}(C_{b_L\mu_L}^{BSM})\text{Re}(C_{b_R\mu_L}^{BSM}) + \text{Re}(C_{b_L\mu_R}^{BSM})\text{Re}(C_{b_R\mu_R}^{BSM})]. \end{aligned} \quad (15)$$

Another relevant clean observable for studying constraints on the NP coefficient is the branching ration  $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$ , which in terms of NP coefficients reads

$$\text{BR}(B_s \rightarrow \mu^+ \mu^-) = \text{BR}(B_s \rightarrow \mu^+ \mu^-)_{\text{SM}} \left| \frac{C_{b_{L-R}\mu_{L-R}}}{C_{b_{L-R}\mu_{L-R}}^{\text{SM}}} \right|^2, \quad (16)$$

<sup>1</sup> It is worth to mention that the analytic formulas given in Eqs. (9), and (10) provide a less precise approximation for the lower bin  $[0.1, 1.1]$ . Nevertheless, it is important to observe that measurements of the lower  $q^2$  interval are associated with a higher level of uncertainty compared to the higher bin.

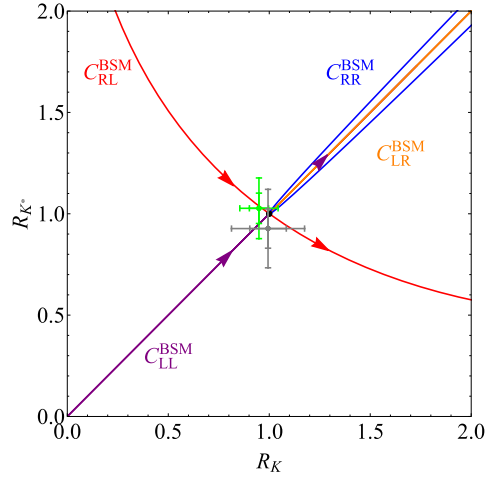


Fig. 1. Quantitative behaviour of  $R_K$  vs  $R_{K^*}$  as in (9)-(10), normalised by their SM value, obtained by switching on one NP coefficient at a time. The experimental values for the  $[1.1, 6]$  GeV<sup>2</sup> bin (in green) and for the  $[0.1, 1.1]$  GeV<sup>2</sup> bin (in grey) are shown with  $1\sigma$  and  $2\sigma$  error bars. The experimental values are compatible with the SM in the 2-sigma range, signalling values of the NP coefficients close to 0.

with  $\text{BR}(B_s \rightarrow \mu^+ \mu^-)_{\text{SM}} = (3.66 \pm 0.14) \times 10^{-9}$ . Following the most recent measurement by the CMS collaboration [31], the experimental value, as reported in [32], for the branching ratio  $\text{BR}(B_s \rightarrow \mu^+ \mu^-)_{\text{exp}}$  is now  $(3.28 \pm 0.26) \times 10^{-9}$ . This measurement is now even closer to the Standard Model prediction. Inserting the numerical values for the SM Wilson coefficients the above becomes

$$\frac{\text{BR}(B_s \rightarrow \mu^+ \mu^-)}{\text{BR}(B_s \rightarrow \mu^+ \mu^-)_{\text{SM}}} = 1 + 0.227[\text{Re}(C_{b_L \mu_L}^{\text{BSM}}) + \text{Re}(C_{b_R \mu_R}^{\text{BSM}}) - \text{Re}(C_{b_L \mu_R}^{\text{BSM}}) - \text{Re}(C_{b_R \mu_L}^{\text{BSM}})] + 0.013[\text{Re}(C_{b_L \mu_L}^{\text{BSM}}) + \text{Re}(C_{b_R \mu_R}^{\text{BSM}}) - \text{Re}(C_{b_L \mu_R}^{\text{BSM}}) - \text{Re}(C_{b_R \mu_L}^{\text{BSM}})]^2 \quad (17)$$

As a preliminary step, we consider the  $2\text{-}\sigma$  interval for the experimental values presented above in order to obtain constraints on the Wilson coefficients. As mentioned earlier the values of  $R_K$  and  $R_{K^*}$  in the lower  $q^2$  interval are associated with a higher level of uncertainty compared to the higher bin. As a result the relevant constraint stems from the  $q^2$  bin  $[1.1, 6]$  GeV<sup>2</sup>, as well as from the  $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$ .

When switching on one real Wilson coefficient at a time, our analytical analysis, at the  $2\text{-}\sigma$  level, yields the following constraints:

$$C_{b_L \mu_L}^{\text{BSM}} \in [-0.53, 0.19], \quad (18)$$

$$C_{b_L \mu_R}^{\text{BSM}} \in [-0.46, 1.51], \quad (19)$$

$$C_{b_R \mu_L}^{\text{BSM}} \in [-0.46, 0.19], \quad (20)$$

$$C_{b_R \mu_R}^{\text{BSM}} \in [-1.51, 0.46]. \quad (21)$$

These intervals provide a first idea of the acceptable ranges for the associated Wilson coefficients, taking into account uncertainties at  $2\text{-}\sigma$  confidence level. We have also considered the imaginary parts of the Wilson coefficients, but we will concentrate on the real parts since these are preferred by the data.

## 2.2. Fit analysis

So far we examined the essential features of the HI observables, laying the groundwork for our expectations in the comparisons with experimental data.

Building upon the analysis presented in [1,25], this section utilises the FLAVIO toolkit [20] to revise the outcomes of the fitting procedure, refining the values of the Wilson NP coefficients to optimally align with the available dataset. In the initial phase of our analysis, we concentrate exclusively on the HI set of observables. Following this, we estimate the influence of the ‘Hadronic Sensitive’ (HS) observables and subsequently incorporate them into a unified global fitting procedure.

In Table 1, we provide a concise overview of the updated outcomes obtained from the fitting procedure, focusing on the scenario where individual Wilson coefficients are turned on one at a time. To ensure a comprehensive update of the findings presented in [1], we extend our analysis to include situations where NP effects are assumed in the electron sector only.

The obtained results align with the analytic investigation presented in section 2.1. The overall deviation from the SM is significantly mitigated with respect to previous results [1,23,25,33–45]. This is to be expected because the recent experimental measurements align well with the SM predictions. In fact, the muon Wilson coefficients are now in agreement with the SM within a  $1.2\sigma$  confidence level. As for the muon case, the electron sector fit shows compatibility with the SM with a significance of  $0.7\sigma$ . For completeness, we show in Table 2 the results of 1-parameter fits in the vector-axial basis.

**Table 1**

Best fits turning on a single operator at a time in the chiral basis, using the ‘hadronic insensitive’ observables ‘HI’, the ‘hadronic sensitive’ observables ‘HS’, or all the observables ‘all’. The full list of observables can be found in Appendix A.

	New physics in the muon sector (Chiral basis)								
	Best-fit			1- $\sigma$ range			$\sqrt{\chi_{SM}^2 - \chi_{best}^2}$		
	HI	HS	all	HI	HS	all	HI	HS	all
$C_{b_L\mu_L}^{BSM}$	-0.15	-1.31	-0.33	-0.05	-1.05	-0.24	1.1	4.1	2.8
$C_{b_L\mu_R}^{BSM}$	0.40	-0.66	-0.25	0.64	-0.47	-0.10	1.2	2.6	1.7
$C_{b_R\mu_L}^{BSM}$	-0.05	0.08	-0.04	0.05	0.19	0.04	0.3	0.5	0.3
$C_{b_R\mu_R}^{BSM}$	-0.38	0.30	0.05	-0.13	0.52	0.20	1.1	1.6	0.2

	New physics in the electron sector (Chiral basis)								
	Best-fit			1- $\sigma$ range			$\sqrt{\chi_{SM}^2 - \chi_{best}^2}$		
	HI	HS	all	HI	HS	all	HI	HS	all
$C_{b_Le_L}^{BSM}$	0.10	0.94	0.14	0.21	1.45	0.25	0.7	1.2	0.9
$C_{b_Le_R}^{BSM}$	-0.17	-2.71	-0.70	1.03	-1.03	-0.11	0.1	1.3	0.6
$C_{b_Re_L}^{BSM}$	0.14	-3.87	0.15	0.25	-2.86	0.26	0.9	1.4	1.0
$C_{b_Re_R}^{BSM}$	-1.22	-3.94	-1.43	-0.59	-2.92	0.08	0.9	1.4	1.1

In summary, when focusing the analysis on the specific subset of HI observables  $R_K$ ,  $R_{K^*}$  and  $BR(B_s \rightarrow \mu^+ \mu^-)$ , the overall results indicate a compatibility with the Standard Model within nearly a  $1\sigma$  level.

Nevertheless, a significant deviation in the muon sector persists, originating from the HS observables. In the electron sector, however, one observes a reduced deviation from the SM which now hovers around  $1\sigma$ . This is to be expected since the majority of HS observables deviating from the SM involve muons.

We finally combine HS and HI observables in a global fit. When considering the single coefficient turned on, in the muon sector, we conclude that the results favour a deviation in the SM in  $C_{b_L\mu_L}^{BSM}$  as well as  $C_{b_L\mu_R}^{BSM}$ , with a larger significance for the left-handed muon coefficient than the right-handed one. Meanwhile, we observe that  $C_{b_R\mu_L}^{BSM}$  and  $C_{b_R\mu_R}^{BSM}$  remain compatible with the SM value, as do all the electronic coefficients.

Then, we proceed with a comprehensive fit involving multiple Wilson coefficients. Given the reduced sensitivity of our observables to electronic coefficients, we exclusively focus on the muonic operators. When turning on four BSM Wilson coefficients at once, the results of the global fit are:

$$\begin{aligned}
C_{b_L\mu_L}^{BSM} &= -0.65 \pm 0.10, \\
C_{b_L\mu_R}^{BSM} &= -0.93 \pm 0.09, \quad \chi_{SM}^2 = 264.99, \\
C_{b_R\mu_L}^{BSM} &= 0.20 \pm 0.15, \quad \tilde{\chi}^2 = 246.77. \\
C_{b_R\mu_R}^{BSM} &= -0.16 \pm 0.28.
\end{aligned}
\quad \rho = \begin{pmatrix} 1 & 0.15 & -0.36 & -0.29 \\ -0.15 & 1 & 0.28 & 0.39 \\ -0.36 & 0.28 & 1 & 0.83 \\ -0.29 & 0.39 & 0.83 & 1 \end{pmatrix}. \quad (22)$$

We report the reduced chi-square:

$$\frac{\chi_{SM}^2}{\# \text{ d.o.f.}} = 0.985, \quad \frac{\tilde{\chi}^2}{\# \text{ d.o.f.}} = 0.931. \quad (23)$$

When comparing the  $\chi^2$  values, we find a deviation from the SM at the  $4.3\sigma$  level. However, upon using the ‘Pull’ value, as defined in [46]:

$$\text{Pull} = \sqrt{\text{CDF}_1^{-1}(\text{CDF}_{N_{par}}(\chi_{SM}^2 - \tilde{\chi}^2))}, \quad (24)$$

**Table 2**

Best fits turning on a single operator at a time in the vector-axial basis, using the ‘hadronic insensitive’ observables ‘HI’, the ‘hadronic sensitive’ observables ‘HS’, or all the observables ‘all’.

New physics in the muon sector (Vector Axial basis)									
	Best-fit			1- $\sigma$ range			$\sqrt{\chi_{SM}^2 - \chi_{best}^2}$		
	HI	HS	all	HI	HS	all	HI	HS	all
$C_{9,\mu}^{BSM}$	-0.12	-0.94	-0.49	-0.24 0.00	-0.82 -1.06	-0.39 -0.59	0.7	4.9	3.6
$C_{10,\mu}^{BSM}$	0.14	-0.23	0.16	0.22 0.06	0.36 0.09	0.23 0.09	1.3	1.2	1.6
$C'_{9,\mu}{}^{BSM}$	-0.01	0.20	-0.03	0.07 -0.09	0.33 0.07	0.05 -0.11	0.9	1.1	0.3
$C'_{10,\mu}{}^{BSM}$	-0.15	0	0.03	-0.26 -0.04	0.08 -0.08	0.08 -0.02	0.1	0	0.3

New physics in the electron sector (Vector Axial basis)									
	Best-fit			1- $\sigma$ range			$\sqrt{\chi_{SM}^2 - \chi_{best}^2}$		
	HI	HS	all	HI	HS	all	HI	HS	all
$C_{9,e}^{BSM}$	0.12	1.01	0.15	0.24 0.00	1.55 0.47	0.26 0.04	0.7	1.4	1.0
$C_{10,e}^{BSM}$	-0.09	-0.79	-0.12	0.01 -0.19	-0.36 -1.21	-0.02 -0.22	0.6	1.2	0.9
$C'_{9,e}{}^{BSM}$	0.15	0.20	0.16	0.27 0.03	0.33 0.07	0.27 0.05	0.9	1.4	1.0
$C'_{10,e}{}^{BSM}$	-0.14		-0.14	-0.04 -0.24		-0.40 -0.24	0.9		1.0

we also take into account the number of parameters switched on. Here,  $CDF_n$  stands for the cumulative distribution function of a  $\chi^2$ -distributed random variable with  $n$  degrees of freedom, and  $N_{par}$  is the number of fitted parameters. Using (24), the discrepancy with the SM is revealed at a slightly reduced significance level of  $3.3\sigma$ .

Interestingly, we observe that the results in (22) agree within a  $\sigma$  from the findings presented in [1], when the global fit was performed already assuming  $R_K = R_{K^*} = 1$  long before the latest experimental values were available and before any other successive investigation.

### 2.3. Leptoquark motivated fit scenarios

To conclude this section, we now analyse the possibility of NP scenarios involving the simultaneous activation of two Wilson coefficients. As we shall review in section 4, LQ are primary examples of these type of models with the scalar  $S_1 = (\bar{3}, 1, 1/3)$  and vector  $U_1 = (3, 1, 2/3)$  LQs being the relevant ones here because they turn on the following NP coefficients:

$$S_1 : C_{b_L\mu_L}^{BSM}, C_{b_L\mu_R}^{BSM}, \quad (25)$$

$$U_1 : C_{b_L\mu_L}^{BSM}, C_{b_R\mu_R}^{BSM}. \quad (26)$$

The associated contour plots are given in Fig. 2, corresponding to the fitted results:

$$S_1 : \begin{cases} C_{b_L\mu_L}^{BSM} = -0.52 \pm 0.12, \\ C_{b_L\mu_R}^{BSM} = -0.64 \pm 0.14, \end{cases} \quad \text{Pull} = 3.3\sigma \quad \rho = \begin{pmatrix} 1 & 0.54 \\ 0.54 & 1 \end{pmatrix}, \quad (27)$$

$$U_1 : \begin{cases} C_{b_L\mu_L}^{BSM} = -0.31 \pm 0.09, \\ C_{b_R\mu_R}^{BSM} = 0.21 \pm 0.05, \end{cases} \quad \text{Pull} = 1.9\sigma \quad \rho = \begin{pmatrix} 1 & -0.47 \\ -0.47 & 1 \end{pmatrix}. \quad (28)$$

We note that the  $S_1$  model is favoured over the  $U_1$  one when we consider both coefficients turned on simultaneously.

### 3. $b \rightarrow c$ flavour sector

So far, we concentrated on Flavor Changing Neutral Currents (FCNC) process  $b \rightarrow s\ell^+\ell^-$ , it is therefore time to tackle the Flavor Changing Charged Currents (FCCC) sector via the decays  $b \rightarrow c\ell^-\nu_\ell$ . For these, the most general low-energy effective Hamiltonian including operators up to dimension six is:

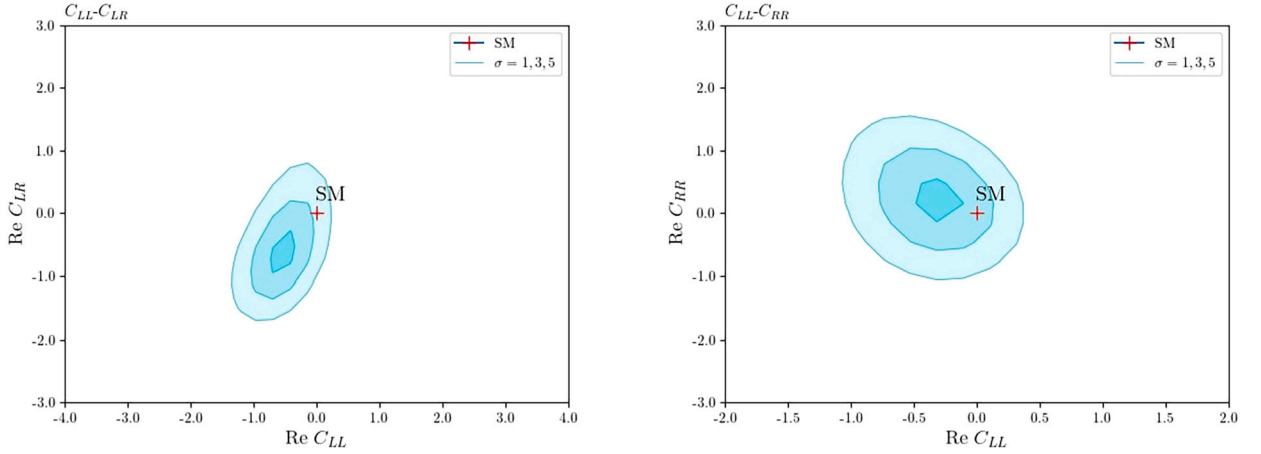


Fig. 2. Contour plots corresponding to the LQ motivated scenarios where we turn on 2 Wilson coefficients at a time, while fixing the remaining coefficients to zero. We use the full list of observables given in Appendix A and present the contour plots corresponding to the 1, 3, and 5- $\sigma$  confidence levels, along with the corresponding SM value.

$$\mathcal{H}_{\text{eff}} = -V_{cb} \frac{\alpha_{\text{em}}}{4\pi v^2} [(1 + C_{V_L})\mathcal{O}_{V_L} + C_{V_R}\mathcal{O}_{V_R} + C_{S_L}\mathcal{O}_{S_L} + C_{S_R}\mathcal{O}_{S_R} + C_T\mathcal{O}_T + \text{h.c.}], \quad (29)$$

with

$$\begin{aligned} \mathcal{O}_{V_L}^\ell &= (\bar{c}_L \gamma^\mu b_L)(\bar{\ell}_L \gamma_\mu \nu_{\ell L}), \\ \mathcal{O}_{V_R}^\ell &= (\bar{c}_R \gamma^\mu b_R)(\bar{\ell}_L \gamma_\mu \nu_{\ell L}), \\ \mathcal{O}_{S_L}^\ell &= (\bar{c}_L b_R)(\bar{\ell}_R \nu_{\ell L}), \\ \mathcal{O}_{S_R}^\ell &= (\bar{c}_R b_L)(\bar{\ell}_R \nu_{\ell L}), \\ \mathcal{O}_T^\ell &= (\bar{c}_R \sigma^{\mu\nu} b_L)(\bar{\ell}_R \sigma_{\mu\nu} \nu_{\ell L}). \end{aligned} \quad (30)$$

The NP contributions, differently from the FCNC case, are directly encoded in the Wilson coefficients (WCs), denoted as  $C_X$ . One can envision including operators accounting for contributions stemming from right-handed neutrinos, [47–49]. These, however, would go beyond the scope of this work which is to keep the level of model building minimal. We now turn to contributions stemming from the tau sector only.

### 3.1. Current status of the $b \rightarrow c$ observables

The semi-tauonic decays of  $B$  mesons, specifically  $B \rightarrow D^{(*)} \tau \bar{\nu}_\tau$ , are crucial processes for probing lepton flavor universality violation (LFUV), complementing the investigation of  $B \rightarrow K^{(*)} \ell \bar{\ell}$  discussed in the preceding section. The ratios

$$R_D = \frac{\text{BR}(B \rightarrow D \tau \bar{\nu}_\tau)}{\text{BR}(B \rightarrow D \ell \bar{\nu}_\ell)}, \quad (31)$$

$$R_{D^*} = \frac{\text{BR}(B \rightarrow D^* \tau \bar{\nu}_\tau)}{\text{BR}(B \rightarrow D^* \ell \bar{\nu}_\ell)}, \quad \ell = e, \mu, \quad (32)$$

have been computed with high precision. This precision is achieved through the remarkable suppression of hadronic uncertainties associated with the strong interaction in the  $B \rightarrow D^{(*)}$  transitions.

The experimental measurements, as detailed in Table 3, consistently reveal deviations from Standard Model predictions. Notably, the characteristic pattern is that the experimental values of  $R_{D^{(*)}}$  exceed the SM predictions, implying a systematic violation of lepton flavor universality (LFU). In Table 3, we provide a concise overview of the current status of independent measurements of  $R_D$  and  $R_{D^*}$  conducted by various collaborations. We also include the latest result on  $R_{D^*}$  from Belle II from the recent presentation in [50]. In addition to the ratios  $R_D$  and  $R_{D^*}$ , significant constraints arise from the  $B_c$  meson branching ratio  $\text{BR}(B_c \rightarrow \tau \nu_\tau)$ . The upper bound on the branching ratio  $\text{BR}(B_c \rightarrow \tau \nu_\tau)$  is given as  $\text{BR}(B_c \rightarrow \tau \nu_\tau) \leq 0.3$  according to [51]. Alternatively, a more conservative limit of 0.6 is suggested in [46], while a more stringent bound of 0.1, based on LEP data, is presented in [52].

In our analysis, we also consider the measurement of the  $D^*$  meson longitudinal polarisation fraction, denoted as  $F_L^{D^*}$ , with a value of  $0.60 \pm 0.08 \pm 0.04$ , as reported by the Belle collaboration [53]. This result is in agreement within approximately  $1.7\sigma$  with the SM prediction.

**Table 3**

Current status of the independent experimental  $R_{D^{(*)}}$  measurements. The first and second errors are statistical and systematic, respectively.

Experiment	$R_{D^*}$	$R_D$	Correlation
BaBar [54,55]	$0.332 \pm 0.024 \pm 0.018$	$0.440 \pm 0.058 \pm 0.042$	-0.27
Belle [56]	$0.293 \pm 0.038 \pm 0.015$	$0.375 \pm 0.064 \pm 0.026$	-0.49
Belle [57,58]	$0.270 \pm 0.035^{+0.028}_{-0.025}$	-	-
Belle [59,60]	$0.283 \pm 0.018 \pm 0.014$	$0.307 \pm 0.037 \pm 0.016$	-0.51
LHCb [61]	$0.281 \pm 0.018 \pm 0.024$	$0.441 \pm 0.060 \pm 0.066$	-0.43
LHCb [62]	$0.257 \pm 0.012 \pm 0.018$	-	-
Belle II [50]	$0.267^{+0.041+0.028}_{-0.039-0.033}$	-	-
World average [4]	$0.284 \pm 0.012$	$0.357 \pm 0.029$	-0.37

### 3.2. Analysis of the $b \rightarrow c$ observables

Using the effective Hamiltonian from Eq. (29), where the NP Wilson coefficients are defined at the renormalisation scale  $\mu = \mu_b = 4.18 \text{ GeV}$ , we employ the updated numerical formulas provided in [63]:

$$\begin{aligned} \frac{R_D}{R_D^{SM}} &= \left| 1 + C_{V_L} + C_{V_R} \right|^2 + 1.01 \left| C_{S_R} + C_{S_L} \right|^2 + 0.84 |C_T|^2 \\ &+ 1.49 \text{Re}[(1 + C_{V_L})(C_{S_R}^* + C_{S_L}^*)] + 1.08 \text{Re}[(1 + C_{V_L} + C_{V_R})C_T^*], \end{aligned} \quad (33)$$

$$\begin{aligned} \frac{R_{D^*}}{R_{D^*}^{SM}} &= \left| 1 + C_{V_L} \right|^2 + \left| C_{V_R} \right|^2 + 0.04 \left| C_{S_L} - C_{S_R} \right|^2 + 16.0 |C_T|^2 \\ &- 0.11 \text{Re} \left[ (1 + C_{V_L} - C_{V_R})(C_{S_L}^* - C_{S_R}^*) \right] - 1.83 \text{Re} \left[ (1 + C_{V_L})C_{V_R}^* \right] \\ &- 5.17 \text{Re} \left[ (1 + C_{V_L})C_T^* \right] + 6.60 \text{Re} \left[ C_{V_R}C_T^* \right], \end{aligned} \quad (34)$$

$$\begin{aligned} \frac{F_L^{D^*}}{F_{L,SM}^{D^*}} &= \left( \frac{R_{D^*}}{R_{D^*}^{SM}} \right)^{-1} \left( \left| 1 + C_{V_L} - C_{V_R} \right|^2 + 0.08 \left| C_{S_L} - C_{S_R} \right|^2 + 6.90 |C_T|^2 \right. \\ &\left. - 0.25 \text{Re} \left[ (1 + C_{V_L} - C_{V_R})(C_{S_L}^* - C_{S_R}^*) \right] - 4.30 \text{Re} \left[ (1 + C_{V_L} - C_{V_R})C_T^* \right] \right), \end{aligned} \quad (35)$$

$$\frac{\text{BR}(B_c^+ \rightarrow \tau^+ \nu_\tau)}{\text{BR}(B_c^+ \rightarrow \tau^+ \nu_\tau)_{SM}} = \left| 1 + C_{V_L} - 4.35(C_{S_L} - C_{S_R}) \right|^2. \quad (36)$$

The SM values that we used for these quantities are given by [4,63]

$$R_D^{SM} = 0.298, \quad R_{D^*}^{SM} = 0.254, \quad F_L^{D^*} = 0.464, \quad \text{BR}(B_c^+ \rightarrow \tau^+ \nu_\tau)_{SM} = 0.022. \quad (37)$$

Using the expressions for  $R_D$  and  $R_{D^*}$ , we illustrate in Fig. 3 the parametric curves corresponding to the single NP Wilson coefficient in the  $R_D - R_{D^*}$  plane, along with the current experimental world average reported in Table 3.

Using the experimental values for the observables in Eqs. (33)–(36) and considering the  $2\text{-}\sigma$  interval, we establish constraints on the Wilson coefficients when turning on a single coefficient at a time. We also explore scenarios motivated by LQ models, defined by the relations  $C_{S_L}(\Lambda_{LQ}) = +4C_T(\Lambda_{LQ})$  and  $C_{S_L}(\Lambda_{LQ}) = -4C_T(\Lambda_{LQ})$ , at the LQ scale  $\Lambda_{LQ} \approx 2 \text{ TeV}$ . After accounting for the renormalisation group (RG) running from  $\Lambda_{LQ}$  to  $\mu_b$ , these relations translate into  $C_{S_L}(\mu_b) = 8.4C_T(\mu_b)$  and  $C_{S_L}(\mu_b) = -8.9C_T(\mu_b)$  respectively [63,64]. The constraints at  $\mu_b$  are:

$$C_{V_L} \in [0.01, 0.10], \quad (38)$$

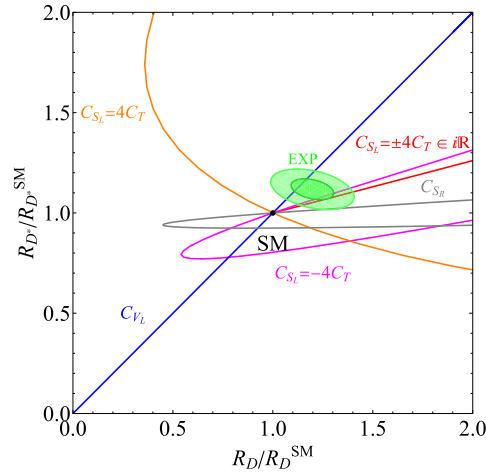
$$C_{S_R} \in [0.20, 0.23], \quad (39)$$

$$C_{S_L} = -8.9C_T \in [0.05, 0.24]. \quad (40)$$

As shown in Fig. 3, the curve corresponding to  $C_{S_L}(\Lambda_{LQ}) = +4C_T(\Lambda_{LQ})$  fails to explain the experimental data at the  $2\text{-}\sigma$  level. It is worth mentioning that, even when considering both  $C_{S_L}$  and  $C_T$  individually, we are unable to reconcile the data. Therefore, we refrained from including the corresponding parametric curves explicitly in the figure.

It is also relevant to consider the scenario where purely imaginary values are allowed for the LQ-motivated coefficients. In this case, we find:





**Fig. 3.** Quantitative behaviour of  $R_D$  vs  $R_D^*$ , normalised by their SM value, obtained by switching on one NP coefficient at a time. We present the current experimental contours at 1 and 2- $\sigma$ , delineated by darker and lighter green regions, respectively. The conditions  $C_{S_L} = \pm 4C_T$  are imposed at 2 TeV and we provide the parametric curves after accounting for the RG running to the low energy scale  $\mu_b$ .

**Table 4**

Best fits at the scale  $\mu_b$  turning on a single operator at a time using the ‘hadronic insensitive’ observables ‘HI’ or all the observables ‘all’. In analogy with the analysis of  $b \rightarrow s$  data, here with HI observables we refer to Eqs. (33)–(36). The full list of observables can be found in Appendix A.

	New physics in the tau sector					
	Best-fit		1- $\sigma$ range		$\sqrt{\chi_{SM}^2 - \chi_{best}^2}$	
	HI	all	HI	all	HI	all
$C_{V_L}$	0.08	0.08	0.09 0.07	0.09 0.07	4.3	4.8
$C_{V_R}$	-0.06	-0.07	-0.04 -0.08	-0.05 -0.09	2.0	2.7
$C_{S_L}$	0.16	0.16	0.20 0.12	0.20 0.12	2.6	2.8
$C_{S_R}$	0.19	0.20	0.22 0.16	0.23 0.17	3.8	4.1
$C_T$	-0.03	-0.03	0.03 -0.09	0.02 -0.08	3.5	4.0

$$|\text{Im}C_{S_L}| = 8.9|\text{Im}C_T| \in [0.31, 0.62], \quad (41)$$

$$|\text{Im}C_{S_L}| = 8.4|\text{Im}C_T| \in [0.30, 0.62]. \quad (42)$$

### 3.3. Fit analysis

We employ the FLAVIO package [20] to perform a fitting analysis of the Wilson coefficients at the scale  $\mu_b$  to match the experimental data.

In order to see how NP scenarios improve the fit, especially when turning on more than one coefficient at a time, we use the ‘Pull’ value defined in Eq. (24), which is expressed in terms of standard deviations  $\sigma$  [46]. Similar to our approach in the  $b \rightarrow s$  sector, our initial analysis focuses solely on the observables defined in Eqs. (33)–(36). Also in this case we refer to these observables as ‘hadronic insensitive’ as they are characterised by a strong suppression of hadronic uncertainties and thus predicted in the SM with an error up to few percent. Subsequently, we extend our analysis to include the effects of all available data summarised in Tables (8 – 10).

We systematically turn on each Wilson coefficient individually, assuming real values. Our results in Table 4 align with those reported in [63] within a 1- $\sigma$  level. We notice a preference for NP coming from a nonzero value of the  $C_{V_L}$  Wilson coefficient with a significance of  $4.3\sigma$ .

Upon allowing the Wilson coefficients to take on complex values, a significant improvement is observed in the fitting for  $C_{V_R}$  and  $C_{S_L}$ , while no enhancement is obtained for  $C_{V_L}$ ,  $C_{S_R}$ , and  $C_T$ . When considering the set of observables given in Eqs. (33)–(36), the best fits along with their corresponding Pull values are:

$$C_{V_R} = 0.01 \pm 0.43i \pm (0.02 + 0.04i), \quad \text{Pull} = 3.9\sigma, \quad (43)$$

$$C_{S_L} = -0.12 \pm 0.63i \pm (0.08 + 0.07i), \quad \text{Pull} = 3.3\sigma. \quad (44)$$

Slightly higher Pull values are achieved when considering the complete set of observables.

### 3.4. LQ motivated fit scenarios

We now proceed to present the results of the fit analysis, considering additional scenarios motivated by the LQ models that we will explore in the next section. We focus exclusively on the set of HI observables specified in Eqs. (33)–(36).

The LQ models along with the combinations of Wilson coefficients not previously considered, evaluated at the LQ scale  $\Lambda_{LQ}$ , are:

$$S_1 : C_{V_L}, C_{S_L} = -4C_T, \quad (45)$$

$$U_1 : C_{V_L}, C_{S_R}, \quad (46)$$

$$R_2 : C_{S_L} = 4C_T. \quad (47)$$

Here we have the following SM quantum numbers for  $R_2 = (3, 2, 7/6)$ .

After accounting for the RG running from  $\Lambda_{LQ}$  to  $\mu_b$ , the relations  $C_{S_L}(\Lambda_{LQ}) = +4C_T(\Lambda_{LQ})$  and  $C_{S_L}(\Lambda_{LQ}) = -4C_T(\Lambda_{LQ})$  become respectively  $C_{S_L}(\mu_b) = 8.4C_T(\mu_b)$  and  $C_{S_L}(\mu_b) = -8.9C_T(\mu_b)$  [63,64]. For these two scenarios we allow Wilson coefficients to assume complex values. In this case, the best fits along with their corresponding Pull values (see Eq. (24)) are:

$$C_{S_L}(\mu_b) = 8.4C_T(\mu_b) = -0.01 \pm 0.04, \quad \text{Pull} = 0.2\sigma, \quad (48)$$

$$C_{S_L}(\mu_b) = 8.4C_T(\mu_b) = \pm 0.53i \pm 0.05i, \quad \text{Pull} = 4.2\sigma, \quad (49)$$

$$C_{S_L}(\mu_b) = 8.4C_T(\mu_b) = -0.07 \pm 0.55i \pm (0.04 + 0.05i), \quad \text{Pull} = 4.0\sigma, \quad (50)$$

$$C_{S_L}(\mu_b) = -8.9C_T(\mu_b) = 0.18 \pm 0.03, \quad \text{Pull} = 4.0\sigma, \quad (51)$$

$$C_{S_L}(\mu_b) = -8.9C_T(\mu_b) = \pm 0.53i \pm 0.05i, \quad \text{Pull} = 4.1\sigma, \quad (52)$$

$$C_{S_L}(\mu_b) = -8.9C_T(\mu_b) = 0.07 \pm 0.44i \pm (0.08 + 0.13i), \quad \text{Pull} = 3.8\sigma. \quad (53)$$

In the case where  $C_{S_L}(\mu_b) = 8.4C_T(\mu_b)$ , the fit indicates a clear preference for complex values. However, in general, the LQs  $S_1$  and  $U_1$  turn on two independent Wilson coefficients simultaneously. While the coefficients  $C_{V_L}$  and  $C_{S_R}$  are preferred to be real, we consider the two cases when the combination  $C_{S_L} = -8.9C_T$  is either real or purely imaginary. We show the contour plots in Fig. 4 and Fig. 5 corresponding to the fit results:

$$S_1 : \begin{cases} C_{V_L} = 0.09 \pm 0.04, \\ C_{S_L} = -8.9C_T = -0.01 \pm 0.09, \end{cases} \quad \text{Pull} = 3.9\sigma \quad \rho = \begin{pmatrix} 1 & -0.93 \\ -0.93 & 1 \end{pmatrix}, \quad (54)$$

$$S_1 : \begin{cases} C_{V_L} = 0.06 \pm 0.03, \\ C_{S_L} = -8.9C_T = (\pm 0.34 \pm 0.13)i, \end{cases} \quad \text{Pull} = 4.0\sigma \quad \rho = \begin{pmatrix} 1 & 0.84 \\ 0.84 & 1 \end{pmatrix}, \quad (55)$$

$$U_1 : \begin{cases} C_{V_L} = 0.07 \pm 0.02, \\ C_{S_R} = 0.06 \pm 0.06, \end{cases} \quad \text{Pull} = 4.0\sigma \quad \rho = \begin{pmatrix} 1 & -0.77 \\ -0.77 & 1 \end{pmatrix}. \quad (56)$$

Differently from the case of  $b \rightarrow s$  the HI observables hint to the presence of new physics in the  $b \rightarrow c$  sector at four sigma level, provided future experimental analyses will avail the current results.

## 4. The new physics landscape

The burning question for present and future experiments in particle physics is: *Where is the scale of new physics?* This overarching goal of the present section is to answer this question by comparing time-honoured extensions of the SM, such as  $Z'$  and LQs models, with  $b \rightarrow s$  and  $b \rightarrow c$  observables.

### 4.1. Shifting up the $Z'$ - boson

$Z'$  bosons have a long history [65–86] and can affect FCNC, with impact for the ratios  $R_{K^{(*)}}$ , via their interactions

$$[g_{bs}(\bar{s}\gamma_\mu P_L b) + \text{h.c.}] + g_{\mu_L}(\bar{\mu}\gamma_\mu P_L \mu), \quad (57)$$

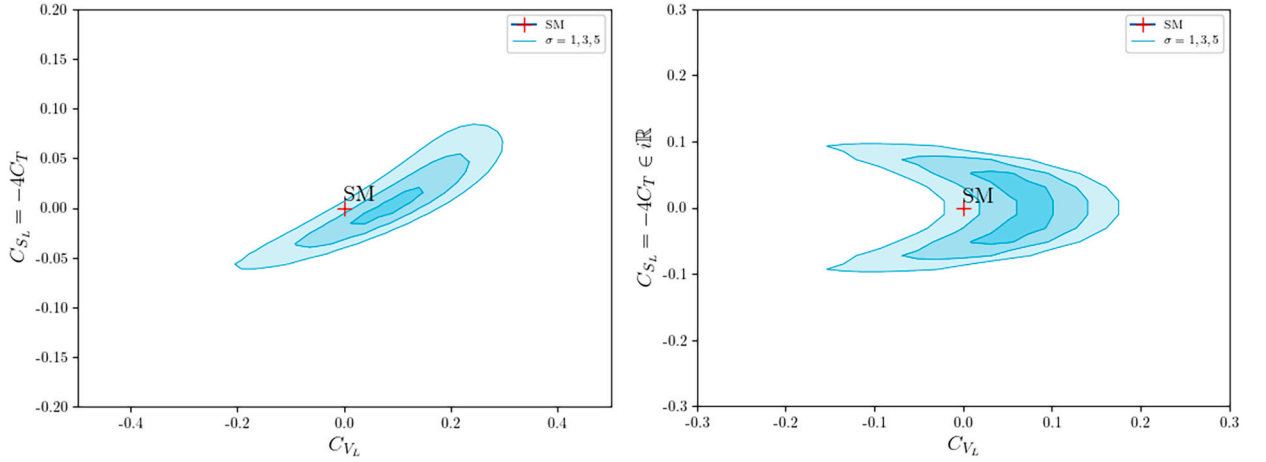


Fig. 4. Contour plots illustrating the behaviour of NP coefficients  $C_{V_L}$ ,  $C_{S_L}$ , and  $C_T$ , considering the theoretical relation  $C_{S_L}(\Lambda_{LQ}) = -4C_T(\Lambda_{LQ})$  predicted by the LQ  $S_1$ . The left panel represents the case where  $C_{S_L} = -4C_T$  is real, while in the right panel it is purely imaginary. We present contour plots corresponding to the 1, 3, and 5- $\sigma$  confidence levels, along with the corresponding SM value.

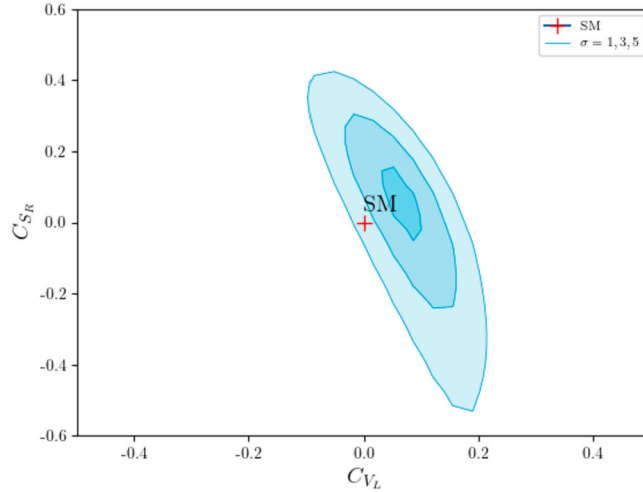


Fig. 5. Contour plot for NP coefficients  $C_{V_L}$  and  $C_{S_R}$ . We present contour plots corresponding to the 1, 3, and 5- $\sigma$  confidence levels, along with the corresponding SM value.

which generates the  $b \rightarrow s\ell^+\ell^-$  transitions. At tree level, a  $Z'$  with the above interactions and mass  $M_{Z'}$  yields

$$C_{b_L\mu_L}^{\text{BSM}} = -\frac{4\pi v^2}{V_{ib}V_{i\bar{s}}^* \alpha_{\text{em}}} \frac{g_{bs}g_{\mu_L}}{M_{Z'}^2}. \quad (58)$$

With the current measurements, this scenario pushes the lower bound for the  $Z'$  mass to about 50 TeV, assuming the product of the couplings to be of order unity for perturbative realizations. For this estimate we used the analytic results in (18) stemming from a direct comparison with the HI observables. We checked that if we consider the global fit with including the HS observables the result does not change.

#### 4.2. Shrinking the leptoquarks landscape

Leptoquarks as bosonic particles carrying both quark and lepton numbers, have been extensively discussed in the literature [38, 87–105] as compelling model building especially when trying to explain possible flavor anomalies. Here we list the relevant LQs, classified by their quantum numbers under the SM gauge group  $(SU(3)_c, SU(2)_L, U(1)_Y)$ , and indicate for each LQ its compatibility with the up-to-date flavor data.

#### 4.2.1. Scalar and vector leptoquarks

We consider three types of scalar LQs ( $S_1, R_2, S_3$ ) and three types of vector LQs ( $U_1, V_2, U_3$ ) which may still be compatible with the new picture emerging from the flavor data. We enforce baryon number conservation to prevent proton decay and further ignore couplings with right-handed neutrinos.

- $S_1 = (\bar{3}, \mathbf{1}, 1/3)$

The interaction of the weak singlet scalar LQ reads:

$$y_{1ij}^{LL} \bar{Q}_L^{i,a} S_1 (i\tau^2)^{ab} L_L^{j,b} + y_{1ij}^{RR} \bar{u}_R^C{}^i S_1 e_R^j + \text{h.c.} . \quad (59)$$

The  $S_1$  interactions can mediate at tree-level charged transitions  $b \rightarrow c\tau\bar{\nu}_\tau$ , with effective couplings:

$$C_{S_L} = -4C_T = -\frac{v^2}{2V_{cb}} \frac{y_{1b\tau}^{LL} (y_{1c\tau}^{RR})^*}{m_{S_1}^2} , \quad (60)$$

$$C_{V_L} = \frac{v^2}{2V_{cb}} \frac{y_{1b\tau}^{LL} (V y_{1c\tau}^{LL*})}{m_{S_1}^2} . \quad (61)$$

With these relations,  $S_1$  can explain the observed anomalies in  $R_{D^{(*)}}$ . Neutral current transitions  $b \rightarrow s\mu\mu$  can be also generated, but only radiatively:

$$C_{b_L\mu_R}^{\text{BSM}} = \frac{m_t^2}{8\pi\alpha_{\text{em}}m_{S_1}^2} (V^* y_{1t\mu}^{LL})_{t\mu} (V y_{1t\mu}^{LL})_{t\mu}^* - \frac{v^2}{16\pi\alpha_{\text{em}}m_{S_1}^2} \frac{(y_{1t\mu}^{LL} \cdot y_{1t\mu}^{\dagger LL})_{bs}}{V_{tb}V_{ts}^*} (y_{1t\mu}^{\dagger LL} \cdot y_{1t\mu}^{LL})_{\mu\mu} , \quad (62)$$

$$C_{b_L\mu_L}^{\text{BSM}} = \frac{m_t^2}{8\pi\alpha_{\text{em}}m_{S_1}^2} (y_{1t\mu}^{RR}) (y_{1t\mu}^{*RR}) \left[ \log\left(\frac{m_{S_1}^2}{m_t^2}\right) - f(x_t) \right] - \quad (63)$$

$$- \frac{v^2}{16\pi\alpha_{\text{em}}m_{S_1}^2} \frac{(y_{1t\mu}^{LL} \cdot y_{1t\mu}^{\dagger LL})_{bs}}{V_{tb}V_{ts}^*} (y_{1t\mu}^{\dagger LL} \cdot y_{1t\mu}^{LL})_{\mu\mu} , \quad (63)$$

where

$$f(x_t) = 1 + \frac{3}{x_t - 1} \left( \frac{\log(x_t)}{x_t - 1} - 1 \right), \quad x_t = \frac{m_t^2}{M_W^2} .$$

Therefore with order one couplings, we find that  $S_1$  can explain both the  $b \rightarrow c$  anomalies and the SM-like results for the  $b \rightarrow s$  processes. The associated mass scale is of the order of a few TeV:

$$1.5 \text{ TeV} \lesssim m_{S_1} \lesssim 3.6 \text{ TeV} . \quad (64)$$

We used the analytic results for the HI observables for both type of transitions. With the upper bound coming from explaining the presence of  $b \rightarrow c$  anomalies and the lower bound from not upsetting the SM-like  $b \rightarrow s$  processes. If, in the future, all anomalies would disappear the upper bound will be lifted.

- $R_2 = (\mathbf{3}, \mathbf{2}, 7/6)$

$R_2$  is a weak doublet with interactions preserving the baryon number that are:

$$-y_{2ij}^{RL} \bar{u}_R^j R_2^a (i\tau^2)^{ab} L_L^{j,b} + y_{2ij}^{LR} \bar{e}_R^i R_2^{a*} Q_L^{j,a} + \text{h.c.} . \quad (65)$$

This model can accommodate the data on  $R_{D^{(*)}}$  by generating at tree level the charged transition  $b \rightarrow c\tau\bar{\nu}_\tau$ , with an effective coefficient:

$$C_{S_L} = 4C_T = \frac{v^2}{2V_{cb}} \frac{y_{2c\tau}^{LR} (y_{2b\tau}^{RL})^*}{m_{R_2}^2} . \quad (66)$$

In fact one can generate, still at tree level, the coefficient

$$C_{b_L\mu_R}^{\text{BSM}} = -\frac{2\pi v^2}{V_{tb}V_{ts}^* \alpha_{\text{em}}} \frac{y_{2s\mu}^{LR} y_{2b\mu}^{LR*}}{m_{R_2}^2} , \quad (67)$$

relevant for the  $b \rightarrow s$  processes. As a result this LQ can accommodate the  $b \rightarrow c$  anomalies while agreeing with the  $b \rightarrow s$  results only for specific structures of couplings such as small couplings to muons than to taus. If, however, we assume all the couplings of order one, the  $b \rightarrow s$  data would require over 20 TeV for the mass of  $R_2$  which would mean that the  $b \rightarrow c$  data are not compatible with this scenario.

- $S_3 = (\bar{3}, 3, 1/3)$

The Lagrangian of the weak triplet is given by:

$$y_{3ij}^L \bar{Q}_L^{i,a} (i\tau^2)^{ab} (\tau^k S_3^k)^{bc} L_L^{j,c} + \text{h.c.} \quad (68)$$

Note that the couplings to diquarks are suppressed in order to guarantee proton stability.  $S_3$  can mediate neutral current transitions  $b \rightarrow s\mu\mu$  at tree-level. After integrating out the LQ, one finds

$$C_{b_L\mu_L}^{\text{BSM}} = \frac{4\pi v^2}{V_{ib} V_{is}^* \alpha_{\text{em}}} \frac{y_{3b\mu}^L (y_{3s\mu}^L)^*}{m_{S_3}^2}. \quad (69)$$

However  $S_3$  cannot address the data  $R_{D^{(*)}}^{\text{exp}} > R_{D^{(*)}}^{\text{SM}}$ , since it generates a negative coefficient  $C_{V_L}$  (after accounting for the data on  $B \rightarrow K^{(*)} \nu \bar{\nu}$  and  $\Delta m_{B_s}$ ) [106]

$$C_{V_L} = -\frac{v^2}{V_{cb}} \frac{(V y_3^L)^{c\tau} y_{3b\tau}^{L*}}{m_{U_3}^2}. \quad (70)$$

Employing the constraint given in Eq. (18), we determine that the minimum threshold for the LQ scale is approximately 50 TeV.

- $U_1 = (3, 1, 2/3)$

$U_1$  does not couple to diquarks and its lepto-quark interactions read:

$$x_{1ij}^{LL} \bar{Q}_L^{i,a} \gamma^\mu U_{1,\mu} L_L^{j,a} + x_{1ij}^{RR} \bar{d}_R^i \gamma^\mu U_{1,\mu} e_R^j + \text{h.c.} \quad (71)$$

The absence of diquark interactions, and thus of proton stability issues, render this type of LQ model appealing.  $U_1$  generates the charged transition  $b \rightarrow c\tau\bar{\nu}_\tau$ , with the contribution

$$C_{V_L} = \frac{v^2}{V_{cb}} \frac{(V x_1^{LL})_{c\tau} x_{1b\tau}^{LL*}}{m_{U_1}^2}, \quad (72)$$

$$C_{S_R} = \frac{v^2}{2V_{cb}} \frac{(V x_1^{LL})_{c\tau} x_{1b\tau}^{RR*}}{m_{U_1}^2},$$

which can accommodate  $R_{D^{(*)}}$ . Neglecting couplings to right-handed neutrinos, one also predicts the following Wilson's coefficients

$$C_{b_L\mu_L}^{\text{BSM}} = -\frac{4\pi v^2}{V_{ib} V_{is}^* \alpha_{\text{em}}} \frac{x_{1s\mu}^{LL} x_{1b\mu}^{LL*}}{m_{U_1}^2}, \quad (73)$$

$$C_{b_R\mu_R}^{\text{BSM}} = -\frac{4\pi v^2}{V_{ib} V_{is}^* \alpha_{\text{em}}} \frac{x_{1s\mu}^{RR} x_{1b\mu}^{RR*}}{m_{U_1}^2}.$$

Switching on the right-handed couplings, denoted as  $x_1^{RR}$ , results in contributions to the Wilson coefficient  $C_{b_R\mu_R}^{\text{BSM}}$ . As mentioned in section 2.3, current  $b \rightarrow s$  data disfavour this Wilson coefficient. Therefore, we set the right-handed couplings to zero  $x_1^{RR} = 0$ . This scenario is referred to as the minimal  $U_1$  model, as described in [107], where only  $C_{b_L\mu_L}^{\text{BSM}}$  and  $C_{V_L}$  are turned on.

In the natural scenario where the couplings  $(x_1^{LL})^{b(c)\tau}$  and  $x_{1b(s)\mu}^{LL}$  are all of the same order of magnitude, the constraint in Eq. (18), from HI data for the  $b \rightarrow s$  sector would require over 50 TeV for the mass of  $U_1$ . However, the  $b \rightarrow c$  data are not compatible with this scenario.

The LQ  $U_1$  can accommodate both the  $b \rightarrow c$  and  $b \rightarrow s$  HI data under specific coupling structures, requiring a certain fine-tuning:

$$\frac{(V x_1^{LL})_{c\tau} x_{1b\tau}^{LL*}}{x_{1s\mu}^{LL} x_{1b\mu}^{LL*}} \sim 30.$$

Additionally, it is worth mentioning that the minimal  $U_1$  scenario fails to account the recent findings from Belle II experiment concerning the branching ratio  $\text{BR}(B \rightarrow K\nu\nu)$ , as detailed in [108].

- $V_2 = (3, 2, 5/6)$

The Lagrangian of the vector weak doublet  $V_2$  is

$$x_{2ij}^{RL} \bar{d}_R^i \gamma^\mu V_{2,\mu}^a (i\tau^2)^{ab} L_L^{j,b} + x_{2ij}^{LR} \bar{Q}_L^{i,a} \gamma^\mu (i\tau^2)^{ab} V_{2,\mu}^b e_R^j + \text{h.c.} \quad (74)$$

These interactions lead to the following contribution to the  $b \rightarrow s\mu\mu$  process

$$C_{b_L\mu_R}^{\text{BSM}} = -\frac{4\pi v^2}{V_{ib} V_{is}^* \alpha_{\text{em}}} \frac{x_{2s\mu}^{LR} x_{2b\mu}^{LR*}}{m_{V_2}^2}, \quad (75)$$

which is compatible with the HI data. The contribution to the  $b \rightarrow c$  sector is generated at tree-level:

**Table 5**

Summary of leptoquark models and the relevant BSM Wilson coefficients they predict at the effective level. Checkmarks denote models successfully explaining clean observables for a particular flavor transition, while crosses signify otherwise. The final column indicates whether the model can explain both sectors simultaneously. Models marked with (✓) reproduce experimental data with small muonic coefficients.

LQ Model	Wilson Coeff.	$b \rightarrow c$	$b \rightarrow s$	$b \rightarrow c + b \rightarrow s$
$S_1$	$C_{b_L\mu_L}^{\text{BSM}}, C_{b_L\mu_R}^{\text{BSM}}, C_{V_L}, C_{S_L} = -4C_T$	✓	✓	✓
$R_2$	$C_{b_L\mu_R}^{\text{BSM}}, C_{S_L} = 4C_T$	✓	✓	(✓)
$S_3$	$C_{b_L\mu_L}^{\text{BSM}}, C_{V_L}$	×	✓	×
$U_1$	$C_{b_L\mu_L}^{\text{BSM}}, C_{V_L}$	✓	✓	(✓)
$V_2$	$C_{b_L\mu_R}^{\text{BSM}}, C_{S_R}$	✓	✓	(✓)
$U_3$	$C_{b_L\mu_L}^{\text{BSM}}, C_{V_L}$	×	✓	×

$$C_{S_R} = -\frac{2v^2}{V_{cb}} \frac{(V_{X_2^{LR}})_{c\tau} X_{2b\tau}^{RL*}}{m_{V_2}^2}. \quad (76)$$

To describe both  $b \rightarrow s$  and  $b \rightarrow c$  HI data, in this case the tuning in the couplings should be:  $\frac{(V_{X_2^{LR}})_{c\tau} X_{2b\tau}^{RL*}}{X_{2s\mu}^{LR} X_{2b\mu}^{LR*}} \sim 50$ . If we assume all the couplings of order one, the constraint given in Eq. (19) obtained from  $b \rightarrow s$  data gives a lower bound of about 30 TeV for the mass of  $V_2$ .

- $U_3 = (\mathbf{3}, \mathbf{3}, 2/3)$

The weak triplet  $U_3$  couples only to left-handed particles via the interaction terms

$$x_{ij}^{LL} \bar{Q}_L^{i,a} \gamma^\mu (\tau^k U_{3,\mu}^k)^{ab} L_L^{j,b} + \text{h.c.} \quad (77)$$

As in the  $U_1$  case, diquark couplings are absent. These interactions lead to the following contribution to the  $b \rightarrow s\mu\mu$  process

$$C_{b_L\mu_L}^{\text{BSM}} = -\frac{4\pi v^2}{V_{ib} V_{is}^* \alpha_{\text{em}}} \frac{x_{3s\mu}^{LL} x_{3b\mu}^{LL*}}{m_{U_3}^2}, \quad (78)$$

which affects the HI data in the  $b \rightarrow s$  sector. However, it cannot explain the experimental results for the  $R_D$  and  $R_{D^*}$ , due to its negative contribution for the  $b \rightarrow c\tau\bar{\nu}_\tau$  process [106]

$$C_{V_L} = -\frac{v^2}{V_{cb}} \frac{(V_{X_3^{LL}})_{c\tau} X_{3b\tau}^{LL*}}{m_{U_3}^2}. \quad (79)$$

Utilizing the constraint presented in Eq. (18), we establish that the minimum threshold for the LQ scale, required to account for the  $b \rightarrow s$  HI data, is approximately 50 TeV.

We summarise our findings in Table 5: all the leptoquark models can be compatible with the SM-like results for the  $b \rightarrow s$  processes. Instead, the  $b \rightarrow c$  sector can be addressed by all LQ models except for the scalar  $S_3$  and the vector  $U_3$ . Intriguingly only the  $S_1$  state can naturally fit the overall results. For  $U_1$ ,  $V_2$  and  $R_2$  a symmetry must be invoked protecting the couplings with muons. In the table we indicate with a (✓) the success in reproducing the experimental data conditioned to requiring new symmetry limits for the muonic coefficients.

## 5. Conclusions

In this work we have analysed the experimental results for the neutral and charged rare  $B$  meson decays, considering the latest measurements of the HI observables by the LHCb and CMS collaborations. We started with a theoretical investigation of the  $\mu/e$  ratios  $R_K$  and  $R_{K^*}$  as well as the process  $B_s \rightarrow \mu^+ \mu^-$ . We studied the impact of complex Wilson coefficients and derived constraints on both their imaginary and real parts. This analysis has been then followed by a comprehensive comparison with experimental results. We find that:

1. The hadronic insensitive observables are currently compatible with the Standard Model prediction within  $1\sigma$  level.
2. When including the hadronic sensitive observables we observe that deviations from the SM persist, with a preference of new physics in the Wilson coefficient  $C_{b_L\mu_L}^{\text{BSM}}$  with a significance of  $4.1\sigma$ .
3. Considering simultaneously all relevant Wilson coefficients and combining both hadronic sensitive and insensitive data into the fit, the deviation from the SM is observed at  $4.3\sigma$  and decreases at the  $3.3\sigma$ , when considering the Pull value.

Moving on to investigate the  $b \rightarrow c$  anomalies, where a violation of leptonic flavor universality is still evident in the latest measurements, we performed an analysis on the complex Wilson coefficients, establishing constraints on both their imaginary and real components. Following this analysis, the comparison with experimental results reveals that:

1. Deviations from the Standard Model predictions persist, showing preference towards the emergence of a non-vanishing real-valued NP induced coefficient  $C_{V_L}$  with a significance of  $4.3\sigma$  when solely focusing on the hadronic insensitive observables.
2. Leptoquark motivated scenarios where two Wilson coefficients are simultaneously turned on were considered, indicating a deviation at  $4\sigma$  when employing the Pull value. This means that these models are well suited as successful extensions of the Standard Model.

Finally, we reviewed different leptoquark models aimed at explaining the deviations from the Standard Model arriving at the conclusions that: The  $S_1$  leptoquark with a mass in the TeV range can naturally explain the SM-like  $b \rightarrow s$  results and the anomalous deviation of the  $b \rightarrow c$  processes. If, however, one is willing to accept unnaturally small couplings to the second generation leptons also the leptoquarks  $U_1$ ,  $V_2$  and  $R_2$  can be employed.

Overall, even though the landscape of new physics theories has considerably shrunk some motivated extensions are still phenomenologically relevant and worth pursuing via experimental searches.

### Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

### Data availability

All data used are publicly available.

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### Appendix A. Observables

In Tables 6–10 we summarise the observables used in addition to the ‘hadronic insensitive’ observables for both  $b \rightarrow s$  and  $b \rightarrow c$  sectors. All bins are treated in the experimental analyses as independent, even if overlapping.

**Table 6**

List of angular observables used in the global fit in addition to the ‘hadronic insensitive’ observables for the  $b \rightarrow s$  sector. The bins highlighted in blue refer to measurements which deviate more than  $2.5\sigma$  from the theoretical prediction.

Angular observables	
Observable	$[q_{\min}^2, q_{\max}^2]$ [GeV <sup>2</sup> ]
LHCb $B^+ \rightarrow K^{*+} \mu\mu$ 2020 [11], $B^0 \rightarrow K^{*0} \mu\mu$ 2020 S [11]	
$\langle F_L \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle S_3 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle S_4 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle S_5 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle S_7 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle S_8 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle S_9 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle A_{FB} \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle P_1 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle P_2 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle P_3 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle P'_4 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle P'_5 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle P'_6 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]
$\langle P'_8 \rangle$	[1.1, 6], [15, 19], [0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19]

(continued on next page)

Table 6 (continued)

Angular observables	
Observable	$[q_{\min}^2, q_{\max}^2]$ [GeV <sup>2</sup> ]
CMS $B \rightarrow K^* \mu \mu$ 2017 [18]	
$\langle P_1 \rangle(B^0 \rightarrow K^* \mu \mu)$	[1, 2], [2, 4.3], [4.3, 6], [16, 19]
$\langle P'_5 \rangle(B^0 \rightarrow K^* \mu \mu)$	[1, 2], [2, 4.3], [4.3, 6], [16, 19]
ATLAS $B \rightarrow K^* \mu \mu$ 2017 [19]	
$\langle F_L \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]
$\langle S_3 \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]
$\langle S_4 \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]
$\langle S_5 \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]
$\langle S_7 \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]
$\langle S_8 \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]
$\langle P_1 \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]
$\langle P'_4 \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]
$\langle P'_5 \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]
$\langle P'_6 \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]
$\langle P'_8 \rangle(B^0 \rightarrow K^* \mu \mu)$	[0.04, 2], [2, 4], [4, 6], [0.04, 4], [1.1, 6], [0.04, 6]

Table 7

List of differential branching ratios used in the global fit in addition to the ‘hadronic insensitive’ observables for the  $b \rightarrow s$  sector. The bins highlighted in blue and red indicate measurements which deviate more than 2.5 and 3.5  $\sigma$  from the theoretical prediction, respectively.

Branching ratios	
Observable	$[q_{\min}^2, q_{\max}^2]$ [GeV <sup>2</sup> ]
LHCb $B^\pm \rightarrow K \mu \mu$ 2014 [5]	
$\frac{d}{dq^2} \text{BR}(B^\pm \rightarrow K \mu \mu)$	[0.1, 0.98], [1.1, 2], [2, 3], [3, 4], [4, 5], [5, 6], [15, 16], [16, 17], [17, 18], [18, 19], [19, 20], [20, 21], [21, 22], [1.1, 6], [15, 22]
LHCb $B^0 \rightarrow K \mu \mu$ 2014 [5]	
$\frac{d}{dq^2} \text{BR}(B^0 \rightarrow K \mu \mu)$	[0.1, 2], [2, 4], [4, 6], [15, 17], [17, 22], [1.1, 6], [15, 22]
LHCb $B^\pm \rightarrow K^* \mu \mu$ 2014 [5]	
$\frac{d}{dq^2} \text{BR}(B^\pm \rightarrow K^* \mu \mu)$	[0.1, 2], [2, 4], [4, 6], [15, 17], [17, 19], [1.1, 6], [15, 19]
LHCb $B^0 \rightarrow K^* \mu \mu$ 2016 [15]	
$\frac{d}{dq^2} \text{BR}(B^0 \rightarrow K^* \mu \mu)$	[0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19], [1.1, 6], [15, 19]
LHCb $B_s \rightarrow \phi \mu \mu$ 2021 [10]	
$\frac{d}{dq^2} \text{BR}(B_s \rightarrow \phi \mu \mu)$	[0.1, 0.98], [1.1, 2.5], [2.5, 4], [4, 6], [15, 17], [17, 19], [1.1, 6], [15, 19]
Babar $B \rightarrow X_s l l$ 2013 [16]	
$\frac{d}{dq^2} \text{BR}(B \rightarrow X_s l l)$	[0.1, 2], [2.0, 4.3], [4.3, 6.8], [1, 6], [14.2, 25]
$\frac{d}{dq^2} \text{BR}(B \rightarrow X_s \mu \mu)$	[0.1, 2], [2.0, 4.3], [4.3, 6.8], [1, 6], [14.2, 25]
$\frac{d}{dq^2} \text{BR}(B \rightarrow X_s e e)$	[0.1, 2], [2.0, 4.3], [4.3, 6.8], [1, 6], [14.2, 25]
Belle $B \rightarrow X_s l l$ 2005 [17]	
$\frac{d}{dq^2} \text{BR}(B \rightarrow X_s l l)$	[0.04, 1], [1, 6], [14.4, 25]



**Table 8**List of branching ratios within the different  $q^2$ -bins employed in the global fit for the  $b \rightarrow c$  sector.

Binned Branching ratios	
Observable	$[q_{\min}^2, q_{\max}^2]$ [GeV <sup>2</sup> ]
Belle $B^+ \rightarrow D\mu\nu_\mu$ 2015 [109]	
$\text{BR}(B^+ \rightarrow D\mu\nu_\mu)$	[0.0, 1.03], [1.03, 2.21], [2.21, 3.39], [3.39, 4.57], [4.57, 5.75], [5.75, 6.93], [6.93, 8.11], [8.11, 9.3], [9.3, 10.48], [10.48, 11.66]
Belle $B^0 \rightarrow D\mu\nu_\mu$ 2015 [109]	
$\text{BR}(B^0 \rightarrow D\mu\nu_\mu)$	[0.0, 1.03], [1.03, 2.21], [2.21, 3.39], [3.39, 4.57], [4.57, 5.75], [5.75, 6.93], [6.93, 8.11], [8.11, 9.3], [9.3, 10.48], [10.48, 11.66]
Belle $B^+ \rightarrow De\nu_e$ 2015 [109]	
$\text{BR}(B^+ \rightarrow De\nu_e)$	[0.0, 1.03], [1.03, 2.21], [2.21, 3.39], [3.39, 4.57], [4.57, 5.75], [5.75, 6.93], [6.93, 8.11], [8.11, 9.3], [9.3, 10.48], [10.48, 11.66]
Belle $B^0 \rightarrow De\nu_e$ 2015 [109]	
$\text{BR}(B^0 \rightarrow De\nu_e)$	[0.0, 1.03], [1.03, 2.21], [2.21, 3.39], [3.39, 4.57], [4.57, 5.75], [5.75, 6.93], [6.93, 8.11], [8.11, 9.3], [9.3, 10.48], [10.48, 11.66]
Babar $B^+ \rightarrow D\ell\nu_\ell$ 2009 [110]	
$\text{BR}(B^+ \rightarrow D\ell\nu_\ell)$	[0.0, 0.97], [0.97, 2.15], [2.15, 3.34], [3.34, 4.52], [4.52, 5.71], [5.71, 6.89], [6.89, 8.07], [8.07, 9.26], [9.26, 10.44], [10.44, 11.63]
Belle $B \rightarrow D^*\mu\nu_\mu$ 2010 [111]	
$\text{BR}_L(B \rightarrow D^*\mu\nu_\mu)$	[0.08, 1.14], [1.14, 2.2], [2.2, 3.26], [3.26, 4.32], [4.32, 5.38], [5.38, 6.44], [6.44, 7.5], [7.5, 8.57], [8.57, 9.63], [9.63, 10.69]
$\text{BR}_T(B \rightarrow D^*\mu\nu_\mu)$	[0.08, 1.14], [1.14, 2.2], [2.2, 3.26], [3.26, 4.32], [4.32, 5.38], [5.38, 6.44], [6.44, 7.5], [7.5, 8.57], [8.57, 9.63], [9.63, 10.69]
Belle $B \rightarrow D^*e\nu_e$ 2010 [111]	
$\text{BR}_L(B \rightarrow D^*e\nu_e)$	[0.08, 1.14], [1.14, 2.2], [2.2, 3.26], [3.26, 4.32], [4.32, 5.38], [5.38, 6.44], [6.44, 7.5], [7.5, 8.57], [8.57, 9.63], [9.63, 10.69]
$\text{BR}_T(B \rightarrow D^*e\nu_e)$	[0.08, 1.14], [1.14, 2.2], [2.2, 3.26], [3.26, 4.32], [4.32, 5.38], [5.38, 6.44], [6.44, 7.5], [7.5, 8.57], [8.57, 9.63], [9.63, 10.69]
Belle $B^0 \rightarrow D^*\ell\nu_\ell$ 2017 [112]	
$\text{BR}(B^0 \rightarrow D^*\ell\nu_\ell)$	[0.0, 1.14], [1.14, 2.2], [2.2, 3.26], [3.26, 4.32], [4.32, 5.38], [5.38, 6.44], [6.44, 7.5], [7.5, 8.57], [8.57, 9.63], [9.63, 10.69]

**Table 9**List of branching ratios within the different bins employed in the global fit for the  $b \rightarrow c$  sector.

Binned Branching ratios	
Observable	$[\cos\theta_\ell^{\min}, \cos\theta_\ell^{\max}]$ [112]
$\text{BR}(B \rightarrow D^*\ell\nu_\ell)$	[-1.0, -0.8], [-0.8, -0.6], [-0.6, -0.4], [-0.4, -0.2], [-0.2, 0.0], [0.2, 0.4], [0.4, 0.6], [0.6, 0.8], [0.8, 1.0]
$[\cos\theta_\ell^{\min}, \cos\theta_\ell^{\max}]$ [112]	
$\text{BR}(B \rightarrow D^*\ell\nu_\ell)$	[-1.0, -0.8], [-0.8, -0.6], [-0.6, -0.4], [-0.4, -0.2], [-0.2, 0.0], [0.2, 0.4], [0.4, 0.6], [0.6, 0.8], [0.8, 1.0]
$[\phi^{\min}, \phi^{\max}]$ [112]	
$\text{BR}(B \rightarrow D^*\ell\nu_\ell)$	[0.0, $\pi/5$ ], [ $\pi/5$ , $2\pi/5$ ], [ $2\pi/5$ , $3\pi/5$ ], [ $3\pi/5$ , $4\pi/5$ ], [ $4\pi/5$ , $\pi$ ], [ $\pi$ , $6\pi/5$ ], [ $6\pi/5$ , $7\pi/5$ ], [ $7\pi/5$ , $8\pi/5$ ], [ $8\pi/5$ , $9\pi/5$ ], [ $9\pi/5$ , $2\pi$ ]

**Table 10**List of total branching ratios used in the global fit for the  $b \rightarrow c$  sector.

Total Branching ratios [113–118]	
Observable	$\text{BR}(B^+ \rightarrow De\nu_e), \text{BR}(B^+ \rightarrow D^*e\nu_e), \text{BR}(B^+ \rightarrow D\mu\nu_\mu)$ $\text{BR}(B^+ \rightarrow D^*\mu\nu_\mu), \text{BR}(B^+ \rightarrow D^*\ell\nu_\ell), \text{BR}(B^0 \rightarrow D^*\ell\nu_\ell)$

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